

Introduction / Special Relativity SR-I ①

≡ Recall: Einsteins Theory of Gravity takes the assumption that the gravitational field is given by a symmetric, non-degenerate bilinear form g defined on spacetime. A coordinate system $\underline{x} \equiv (x^0, x^1, x^2, x^3)$ given on spacetime determines the components $g_{ij}(\underline{x})$, a symmetric non-deg matrix at each \underline{x} . This determines the differential $ds \equiv$ arclength

$$ds^2 = \pm g_{ij} dx^i dx^j.$$

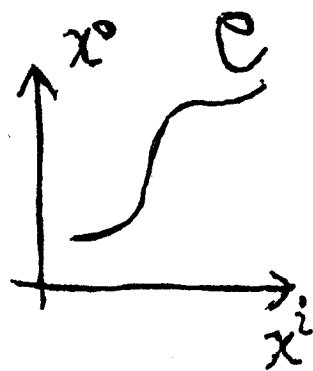
• Given a curve $\underline{x}(\xi)$ on spacetime, the g -length of the curve is given by

$$ds = \sqrt{\pm g_{ij} dx^i dx^j} = \sqrt{\pm g_{ij} \dot{x}^i \dot{x}^j} d\xi$$

since

$$dx^i = \dot{x}^i d\xi.$$

$$\Rightarrow \Delta s = \int_{\xi_1}^{\xi_2} \sqrt{\pm g_{ij} \dot{x}^i \dot{x}^j} d\xi$$



(2)

Assumption ①: $ds = cd\tau$ where $d\tau$ is the proper time change (aging time) for an observer traversing \mathcal{C} .

E.g., if $[x^i] = \text{Meters}$, $[t] = \text{seconds}$, $[x^0 = ct] = \text{meters}$
By taking $ds = cd\tau$, $x^0 = ct$, x^0 and ds have units of meters \Rightarrow space & time have dim. of length.

Assumption ②: ^(timelike) Paths of minimal or critical length \equiv geodesics of g are the freefall paths

Assumption ③: Non-rotating frames are \parallel -transported by connection for g along free-fall paths. (Diff. Geom.)

\equiv
Assumption of Special Relativity: \exists a global coordinate system \underline{x} such that

$$g_{ij} = \begin{bmatrix} -1 & & \\ & 1 & \\ & & 1 \end{bmatrix} \equiv \eta_{ij}$$

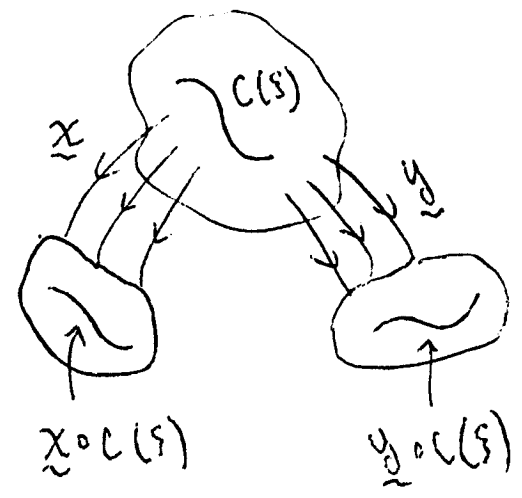
everywhere. \Rightarrow "spacetime is flat"

$\underline{x} \equiv$ Lorentz frame is orthonormal frame for g .

Q: how do g_{ij} transform under a change of coordinates $y^i = y^i(x)$?

Ans: Let $C(s)$ be a curve in spacetime. In x -coords: $x^i(s) \equiv x^i \circ C(s)$

y -coords: $y^\alpha(s) = y^\alpha \circ C(s)$



We want:

$$\int_{\xi_1}^{\xi_2} \sqrt{\pm g_{ij} \dot{x}^i \dot{x}^j} d\xi = \int_{\xi_1}^{\xi_2} \sqrt{\pm \bar{g}_{\alpha\beta} \dot{y}^\alpha \dot{y}^\beta} d\xi$$

so it suffices to make

$$g_{ij} \dot{x}^i \dot{x}^j = \bar{g}_{\alpha\beta} \dot{y}^\alpha \dot{y}^\beta$$

But $y^\alpha = y^\alpha(x) \Rightarrow \dot{y}^\alpha = \frac{\partial y^\alpha}{\partial x^i} \dot{x}^i$

$$\Rightarrow g_{ij} \dot{x}^i \dot{x}^j = \bar{g}_{\alpha\beta} \dot{y}^\alpha \dot{y}^\beta = \bar{g}_{\alpha\beta} \frac{\partial y^\alpha}{\partial x^i} \frac{\partial y^\beta}{\partial x^j} \dot{x}^i \dot{x}^j$$

$$\Rightarrow g_{ij} = \bar{g}_{\alpha\beta} \frac{\partial y^\alpha}{\partial x^i} \frac{\partial y^\beta}{\partial x^j}$$

2. ~~Local~~ Spacetime Manifold: Precisely, we assume Spacetime is a 4-d Manifold of Events which can be covered locally with coordinate charts $\underline{x}: \mathcal{U} \rightarrow \mathbb{R}^4$ st charts are smooth on the overlap:

$\underline{x} \circ \underline{y}^{-1}$ smooth, 1-1, onto, smooth inverse.



The metric g is given, & transforms by

$$g_{ij} = \bar{g}_{\alpha\beta} \frac{\partial y^\alpha}{\partial x^i} \frac{\partial y^\beta}{\partial x^j}$$

• ~~Tangent Space~~ $T\mathcal{M}^4$: Suppose we wish to diff a function $f(x)$ along a curve $c(s) \in \mathcal{M}^4$. We can do this in a coord.

System \underline{x} :

$$\frac{d}{ds} \underline{x}_0$$

• Tangent Space: We'd like a tangent vector to a curve $c(s)$ to be " $X = \frac{dc}{ds}$ ", makes no sense. However, all we use " $\frac{dc}{ds}$ " for is to differentiate functions along the curve

(5)

Let $f(p)$ denote scalar fn

$$f: M^4 \rightarrow \mathbb{R}$$

Then deriv. of f along c :

$\frac{d}{ds} f(c(s))$ is well defined.

Let $f \circ \tilde{x}^{-1}: \mathbb{R}^4 \rightarrow \mathbb{R} \equiv$ "the fn f written in \tilde{x} -coords"

$f \circ \tilde{y}^{-1}: \mathbb{R}^4 \rightarrow \mathbb{R} \equiv$ "the fn f written in \tilde{y} -coords"

Then

curve in \tilde{x} -coords

$$\begin{aligned} \frac{df}{ds}(c(s)) &\equiv \frac{d}{ds} (f \circ \tilde{x}^{-1})(\tilde{x} \circ c(s)) = \frac{\partial f}{\partial x^i} \dot{x}^i = \dot{x}^i \frac{\partial}{\partial x^i} f \\ &= \frac{d}{ds} (f \circ \tilde{y}^{-1})(\tilde{y} \circ c(s)) = \frac{\partial f}{\partial y^a} \dot{y}^a = \dot{y}^a \frac{\partial}{\partial y^a} f \end{aligned}$$

Defn: The tangent vector to curve c is the ⁽⁶⁾
differential operator $X = \dot{x}^i \frac{\partial}{\partial x^i} = \dot{y}^\alpha \frac{\partial}{\partial y^\alpha}$

where $\dot{x}^i \frac{\partial y^\alpha}{\partial x^i} = \dot{y}^\alpha$

This implies that (FIP)

$$\frac{\partial}{\partial x^i} = \frac{\partial y^\alpha}{\partial x^i} \frac{\partial}{\partial y^\alpha},$$

An equality at each pt p

which give the transformation laws for vectors.
We think of $\left\{ \frac{\partial}{\partial x^0}, \dots, \frac{\partial}{\partial x^3} \right\}$ as the x -coordinate basis for the tangent space $T_p M^4$ at pt p .

Thm: $T_p M^4$ is a vector space of dimension 4 (FIP)

• Defn: if $X = a^i \frac{\partial}{\partial x^i}$, $Y = b^j \frac{\partial}{\partial x^j}$, then

$$\langle X, Y \rangle = g_{ij} a^i b^j \text{ so that}$$

$$|X| = |g_{ij} a^i b^j|^{1/2}$$

$$\Delta s = \int_{\xi_1}^{\xi_2} |g_{ij} \dot{x}^i \dot{x}^j|^{1/2} d\xi = \int_{\xi_1}^{\xi_2} \left| \frac{dc}{d\xi} \right| d\xi$$

"vector tangent to curve $c(\xi)$ "

⑦
 [2] SPECIAL RELATIVITY: Assume $g_{ij} = \eta_{ij} = \text{diag}(-1, 1, 1, 1)$
 in x -coordinates. E.g., if we assume

$$ds^2 = \eta_{ij} dx^i dx^j = -dx^0{}^2 + dx^1{}^2 + dx^2{}^2 + dx^3{}^2$$

gives the metric in meters, then $x^0 = ct$,
 t in seconds, $s = c\tau$, τ proper time in sec's.

\Rightarrow proper time change in second for observer
 traversing $C(s)$ is

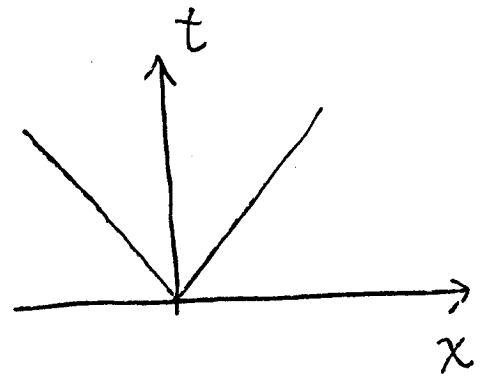
$$\Delta\tau = \frac{1}{c} \int_{\xi_1}^{\xi_2} \sqrt{-\dot{x}^0{}^2 + \dot{x}^1{}^2 + \dots + \dot{x}^3{}^2} d\xi$$

Defn: a vector $X = a^i \frac{\partial}{\partial x^i}$ is

(1) timelike if $\langle X, X \rangle < 0$

(2) lightlike if $\langle X, X \rangle = 0$

(3) spacelike if $\langle X, X \rangle > 0$.



Assumption: The tangent vector to the "world line"
 on curve associated with any particle satisfies
 $\frac{ds}{d\xi}$ is timelike
 "All particles move with speed $< c$ "

Assumption ①: All particles move with speed $< c$ in the \underline{x} -coordinates

Precisely: $\underline{x}(t)$ the world line of a particle

$\Rightarrow \dot{x}^i \frac{\partial}{\partial x^i}$ is timelike. FIP

Assumption ②: The tangent vector to light rays is lightlike. \Leftrightarrow "light rays travel w. speed c "

~~For each O-N frame with points in space can identify vectors~~

We can use the O-N frame \underline{x} to identify vectors in $T_0 M^4$ with points in the space: "identify components with pts in the spacetime"

$$X \equiv X^i \frac{\partial}{\partial x^i} \Big|_0 \leftrightarrow P = \Phi(X): x^i(P) = X^i$$

Under this identification, we can interpret

$$ds(X) \equiv \sqrt{|-(X^0)^2 + \dots + (X^3)^2|}$$

as follows

This is the exponential map FIP

ASSUMPTION (5)

(86)

X timelike: $ds(x) = c d\tau$ is the proper time change between events $P_0 = \Phi(0)$ and $P_1 = \Phi(x)$ as measured by observer moving with velocity vector X ; i.e., if

$$\tilde{x} \circ c(\xi) \equiv \tilde{x}(\xi) \text{ satisfies } \dot{\tilde{x}}(\xi) = X, \tilde{x}(0) = 0$$

then $x^i(\xi) = X^i \xi \Rightarrow$

$$c \Delta\tau = \int_0^1 \sqrt{(X^0)^2 + (X^1)^2 + \dots + (X^3)^2} d\xi = ds(x)$$

Note: if $d\xi \equiv ds$, (arclength param. of $c(\xi)$), then

$$ds = |X| ds \Rightarrow |X| = 1.$$

[i.e., $ds = g_{ij} \dot{x}^i \dot{x}^j d\xi \Rightarrow \langle X, X \rangle = 1$ iff $ds = d\xi$]

in which case we call X the 4-vel. of observer.

~~X spacelike~~

X spacelike: $ds(x)$ is the length in meters of a rod as measured by observer moving in a frame in which P_0 and P_1 occur at same time. E.g., in \underline{x} -frame,

$X^0 = 0 \Rightarrow ds(x) = \sqrt{(x^1)^2 + (x^2)^2 + (x^3)^2}$ is pos def metric giving Euclidean lengths.

• Property of Flat Space: ① You can identify vectors in $T_p M$ with $T_q M$ by:

$$g_{ij} = \eta_{ij} \text{ in coords } \underline{x}^i$$

$$\Leftrightarrow \underline{X}_p = a^i \frac{\partial}{\partial x^i} \Big|_p \leftrightarrow \underline{X}_q = a^i \frac{\partial}{\partial x^i} \Big|_q$$

"vectors with same components at different pts are said to be // translations"

② You can identify $T_{\underline{x}} M$ with M by:

$$\underline{X} = a^i \frac{\partial}{\partial x^i} \Big|_{\underline{x}=0} \leftrightarrow \underline{x}^i = a^i \in \underline{x} \text{ coords of a pt in } M$$

Define: $\underline{X} \in T_p M$, $\mathbb{I}(\underline{X}) = q \in M : \underline{x}^i(q) - \underline{x}^i(p) = \underline{X}^i$

Example: 1-d $g_{ij} = \begin{bmatrix} -1 & 0 \\ 0 & 1 \end{bmatrix}$

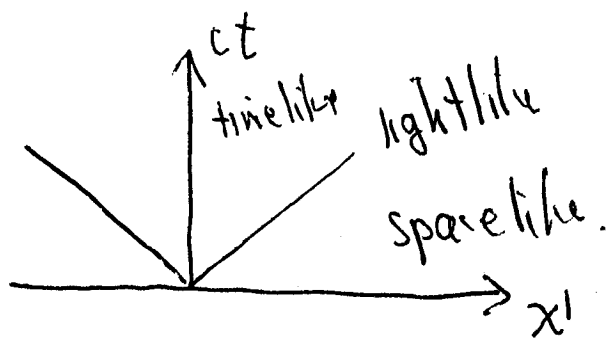
Check: $\left\{ \frac{\partial}{\partial x^0}, \frac{\partial}{\partial x^1} \right\}$ form an o-n basis

$$\left\langle \frac{\partial}{\partial x^0}, \frac{\partial}{\partial x^0} \right\rangle = g_{ij} e_0^i e_0^j = -1$$

$$\left\langle \frac{\partial}{\partial x^1}, \frac{\partial}{\partial x^1} \right\rangle = +1$$

$$\left\langle \frac{\partial}{\partial x^0}, \frac{\partial}{\partial x^1} \right\rangle = 0$$

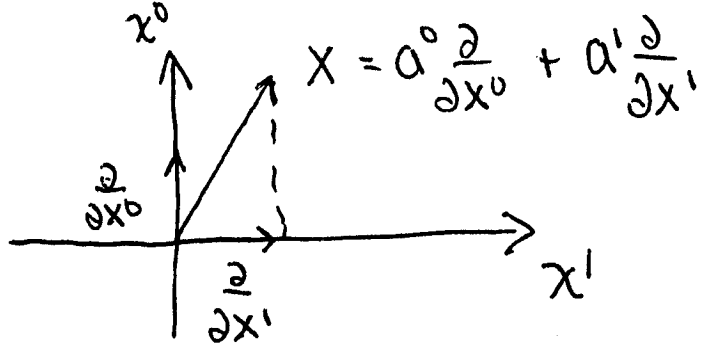
check: $\frac{\partial}{\partial x^0} \pm \frac{\partial}{\partial x^1}$ lightlike $[1, \pm 1] \begin{bmatrix} -1 & 0 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 \\ \pm 1 \end{bmatrix} = 0$



Q: what are the other o-n frames, and how are they related to x -coordinates?

Ans: Let X be vector $X = a^0 \frac{\partial}{\partial x^0} + a^1 \frac{\partial}{\partial x^1}$

Since metric everywhere the same, we can identify components with pts in the space:

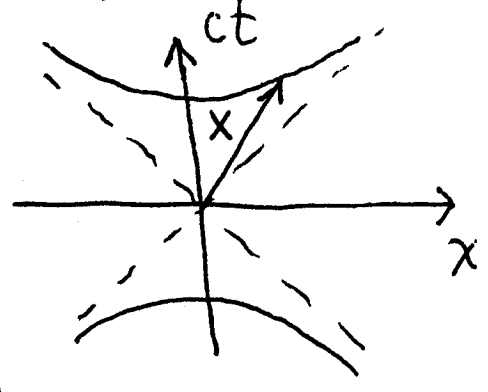


I.e. set $a^0 = ct$, $a^1 = x$, $X = ct \frac{\partial}{\partial x^0} + x \frac{\partial}{\partial x^1}$

Let X be timelike, unit length

$$\langle X, X \rangle = (ct, x) \begin{bmatrix} -1 & 0 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} ct \\ x \end{bmatrix} = -ct^2 + x^2 = -1$$

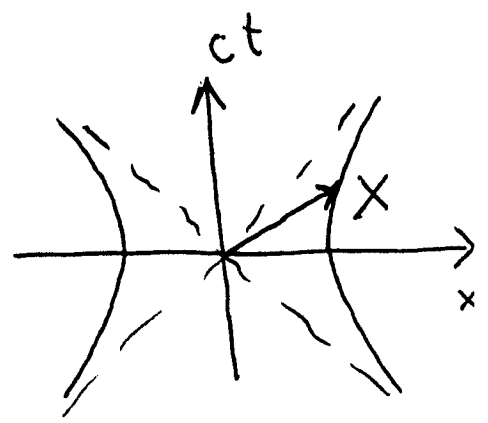
$\Rightarrow X$ lies on unit hyperbola,



Let X be spacelike, unit length

$$\langle X, X \rangle = -ct^2 + x^2 = 1$$

$\Rightarrow X$ lies on unit hyperbola,

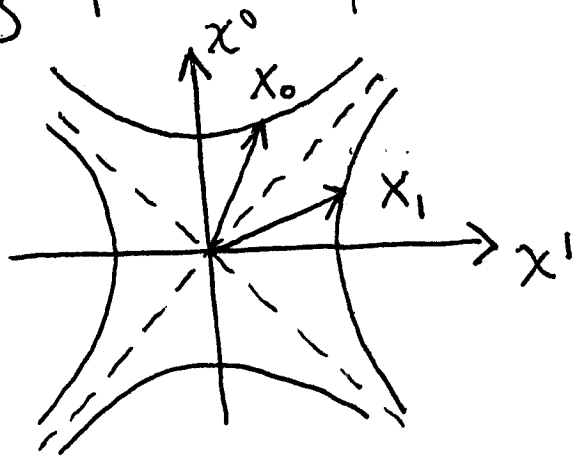


Thus: if $\langle X_0, X_1 \rangle = 0$, $\langle X_0, X_0 \rangle = -1$, $\langle X_1, X_1 \rangle = 1$

$$\langle X_0, X_1 \rangle = (ct_0, x_0) \begin{bmatrix} -1 & 0 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} ct_1 \\ x_1 \end{bmatrix} = (-ct_0, x_0) \cdot (ct_1, x_1)$$

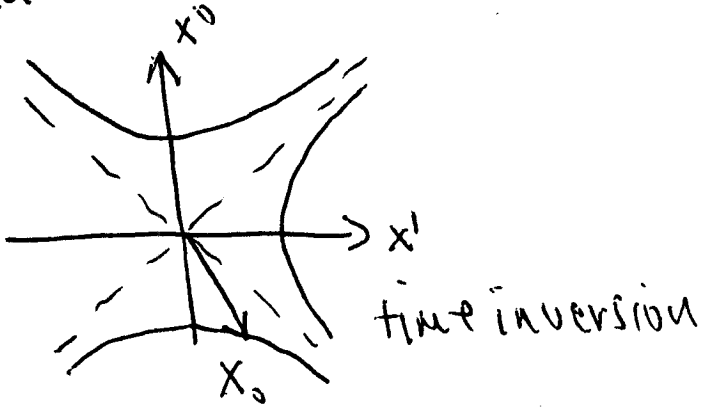
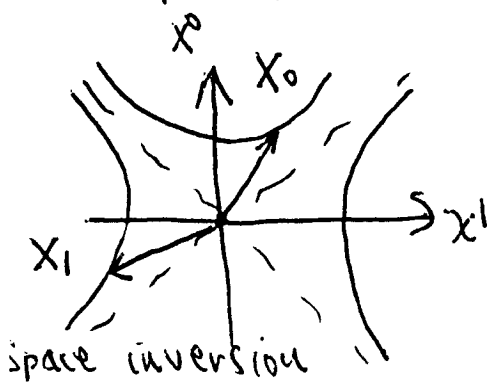
$\Rightarrow "(ct_1, x_1) \parallel \pm (x_0, ct_0) \Rightarrow X_1$ is the reflection of X_0 in line $x = \pm ct$

If we assume X_0 timelike, positive direction and $\{X_0, X_1\}$ positively oriented, then



Note ① cannot get to neg oriented frame thru cont transformations

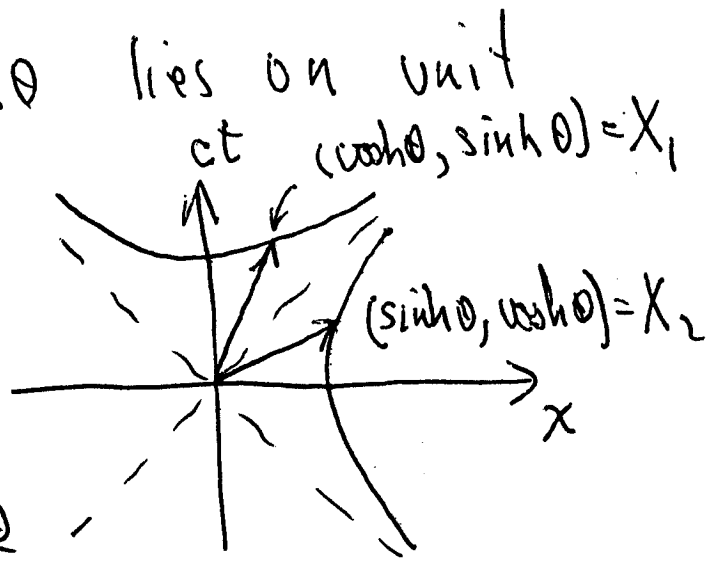
② Cannot get to time inverted or space inverted thru cont sequ. of tran's.



Note: $\cosh^2 \theta - \sinh^2 \theta = 1$

$\Rightarrow ct = \cosh \theta, x = \sinh \theta$ lies on unit hyperbola. Our notation

(a, b) $a \equiv x^0$ -coord
 $b \equiv x^1$ -coord



$\Rightarrow X_0 = \cosh \theta \frac{\partial}{\partial x^0} + \sinh \theta \frac{\partial}{\partial x^1}$

$X_1 = \sinh \theta \frac{\partial}{\partial x^0} + \cosh \theta \frac{\partial}{\partial x^1}$

gives all pos oriented, time oriented, o.n. frames, $-\infty < \theta < \infty$.

Note: ① $\forall X_0$ with $|X_0| = 1$, you can complete it to an o.n. frame. If further X_0 is pos-time directed, then you can complete it uniquely to frame (X_0, X_1) with X_1 pos. space directed.

② All vectors can be completed to o.n. frame except lightlike vectors, which are \perp to themselves.

(3) For any X , $|X| \neq 0$, you can define the orthogonal projection onto X : (15)

$$\text{Proj}_X Y = \frac{\langle X, Y \rangle}{\langle X, X \rangle} X \quad (\text{FIP})$$

↑ keep \pm sign from metric

Using this, you can define the Gram-Schmidt process to construct an O-N basis in \mathbb{R}^4 from any 4 linearly indept non-light like vectors.

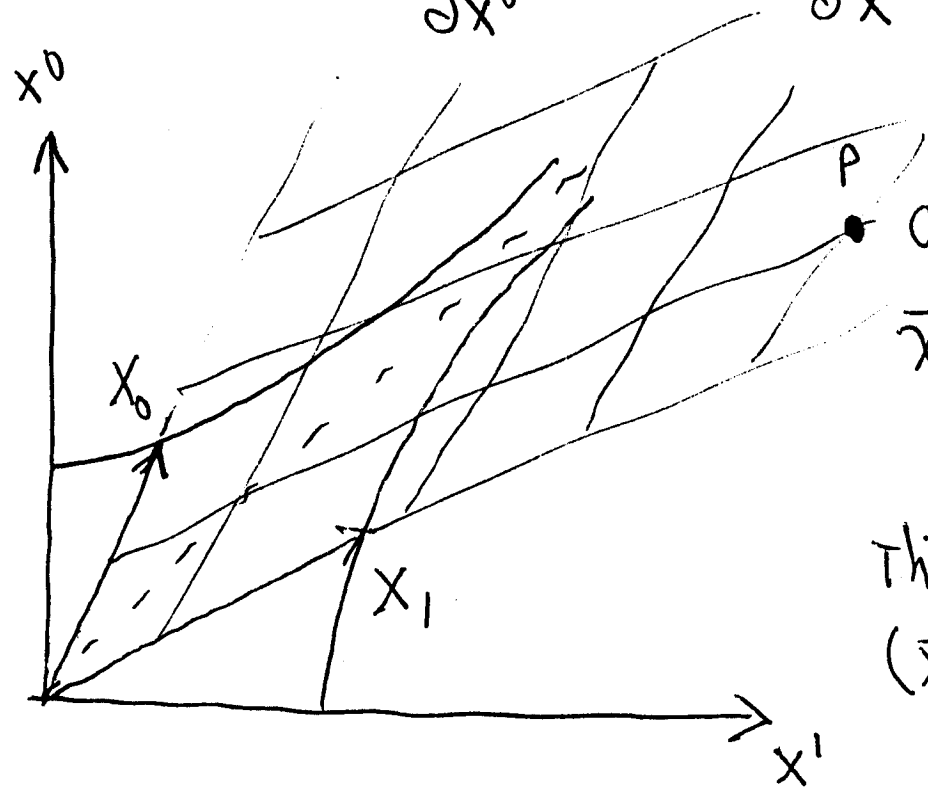
- Lorentz Transformations: Given $g_{ij} = \eta_{ij}$ in $\underline{x} = (x^0, x^1)$ coordinates. Construct another O-N frame for each

$$X_0 = \cosh\theta \frac{\partial}{\partial x^0} + \sinh\theta \frac{\partial}{\partial x^1}$$

$$X_1 = \sinh\theta \frac{\partial}{\partial x^0} + \cosh\theta \frac{\partial}{\partial x^1}$$

I.e., translate these vectors to each point of spacetime ("translation in a flat spacetime") and choose $\{X_0, X_1\}$ to be the coord. basis vectors for a new coord system (\bar{x}^0, \bar{x}^1) on spacetime as follows: We need

$$X_0 = \frac{\partial}{\partial \bar{x}^0}, \quad X_1 = \frac{\partial}{\partial \bar{x}^1}$$



$aX_0 + bX_1$
 $\bar{x}^0(P) = a, \bar{x}^1(P) = b$
 ↑
 This defines the (\bar{x}^0, \bar{x}^1) coord system.

clearly, $\frac{\partial}{\partial \bar{x}^0} = X_0$, $\frac{\partial}{\partial \bar{x}^1} = X_1$, FIP, and

thus $\bar{g}_{ij} = \eta_{ij}$ because $\{X_0, X_1\}$ is an o-n basis at each point of spacetime.

Q: How is the (x^0, x^1) coord. system related to the (\bar{x}^0, \bar{x}^1) coord system?

$$\text{Ans: } x^0 \frac{\partial}{\partial x^0} + x^1 \frac{\partial}{\partial x^1} = \bar{x}^0 \frac{\partial}{\partial \bar{x}^0} + \bar{x}^1 \frac{\partial}{\partial \bar{x}^1}$$

\Leftrightarrow "coord's name same point p"

$$\text{But: } \frac{\partial}{\partial \bar{x}^0} = \cosh \theta \frac{\partial}{\partial x^0} + \sinh \theta \frac{\partial}{\partial x^1}$$

$$\frac{\partial}{\partial \bar{x}^1} = \sinh \theta \frac{\partial}{\partial x^0} + \cosh \theta \frac{\partial}{\partial x^1}$$

$$\Rightarrow x^0 \frac{\partial}{\partial x^0} + x^1 \frac{\partial}{\partial x^1} = (\bar{x}^0 \cosh \theta + \bar{x}^1 \sinh \theta) \frac{\partial}{\partial x^0} + (\bar{x}^0 \sinh \theta + \bar{x}^1 \cosh \theta) \frac{\partial}{\partial x^1}$$

$$\Leftrightarrow \begin{bmatrix} x^0 \\ x^1 \end{bmatrix} = \begin{bmatrix} \cosh \theta & \sinh \theta \\ \sinh \theta & \cosh \theta \end{bmatrix} \begin{bmatrix} \bar{x}^0 \\ \bar{x}^1 \end{bmatrix}$$


Theorem: The positively oriented, time oriented, homogeneous Lorentz transformations are given by $\underline{x} = L(\theta) \underline{\bar{x}}$, where $L(\theta)$ is given by

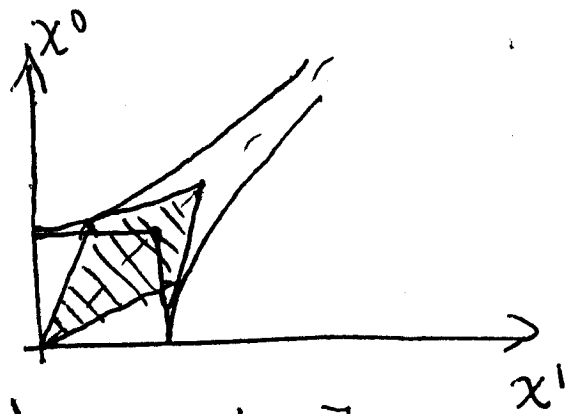
$$\begin{bmatrix} x^0 \\ x^1 \end{bmatrix} = L(\theta) \begin{bmatrix} \bar{x}^0 \\ \bar{x}^1 \end{bmatrix} = \begin{bmatrix} \cosh \theta & \sinh \theta \\ \sinh \theta & \cosh \theta \end{bmatrix} \begin{bmatrix} \bar{x}^0 \\ \bar{x}^1 \end{bmatrix}$$

for $-\infty < \theta < \infty$.

Note ①: $\det L(\theta) = \cosh^2 \theta - \sinh^2 \theta = 1 \Rightarrow$

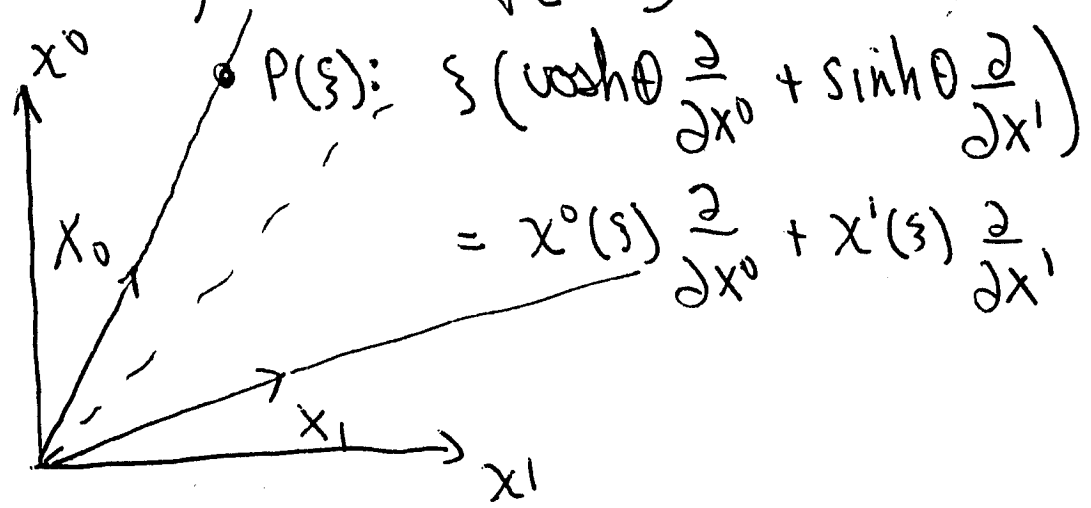
Lorentz transformations preserve the coordinate volume:

i.e., Vol of  = 1.



Note ②: $L(\theta)^{-1} = L(-\theta) = \begin{bmatrix} \cosh \theta & -\sinh \theta \\ -\sinh \theta & \cosh \theta \end{bmatrix}$

Note (2) we can rewrite in terms of the velocity of the \bar{x} -frame as observed in the x -frame by writing $\begin{Bmatrix} \cosh\theta \\ \sinh\theta \end{Bmatrix}$ as a fn of v :



Conclude: $P(s)$ parameterizes the \bar{x}^0 -axis

$$\Leftrightarrow \frac{dx^1}{dx^0} = \frac{dx^1/ds}{dx^0/ds} = \frac{\sinh\theta}{\cosh\theta}$$

$$\begin{cases} x^0(s) = s \cosh\theta \\ x^1(s) = s \sinh\theta \end{cases}$$

$$\therefore \frac{dx^1}{dx^0} = \frac{1}{c} \frac{dx^1}{dt} = \frac{1}{c} v = \frac{\sinh\theta}{\cosh\theta} = \tanh\theta$$

$$1 - \tanh^2\theta = \operatorname{sech}^2\theta = \frac{1}{\cosh^2\theta} \Rightarrow \cosh^2\theta = \frac{1}{1 - (v/c)^2}$$

$$-1 + \tanh^2\theta = \operatorname{coth}^2\theta = \frac{1}{\sinh^2\theta} \Rightarrow \sinh^2\theta = \frac{(v/c)^2}{1 - (v/c)^2}$$

⇒ Lorentz Transformation

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$$\begin{bmatrix} x^0 \\ x^1 \end{bmatrix} = \begin{bmatrix} \frac{1}{\sqrt{1-(v/c)^2}} & \frac{v/c}{\sqrt{1-(v/c)^2}} \\ \frac{v/c}{\sqrt{1-(v/c)^2}} & \frac{1}{\sqrt{1-(v/c)^2}} \end{bmatrix} \begin{bmatrix} \bar{x}^0 \\ \bar{x}^1 \end{bmatrix}$$

gives L-trans, where the bar frame moves with vel. v rel to unbarred frame.

$$\text{Here, } \sqrt{1-(\frac{v}{c})^2} = 1 - \frac{1}{2}(\frac{v}{c})^2 + O(\frac{v}{c})^4$$

$$\frac{1}{\sqrt{1-(\frac{v}{c})^2}} = \frac{1}{1+\{\}} = 1 - \{\} + O(\{\})^2 = 1 + \frac{1}{2}(\frac{v}{c})^2 + O(\frac{v}{c})^4$$

$$\Rightarrow \frac{1}{\sqrt{1-(\frac{v}{c})^2}} = 1 + \frac{1}{2}(\frac{v}{c})^2 + O(\frac{v}{c})^4$$

$$\Rightarrow L(\theta) = Id + O(\frac{v}{c})$$

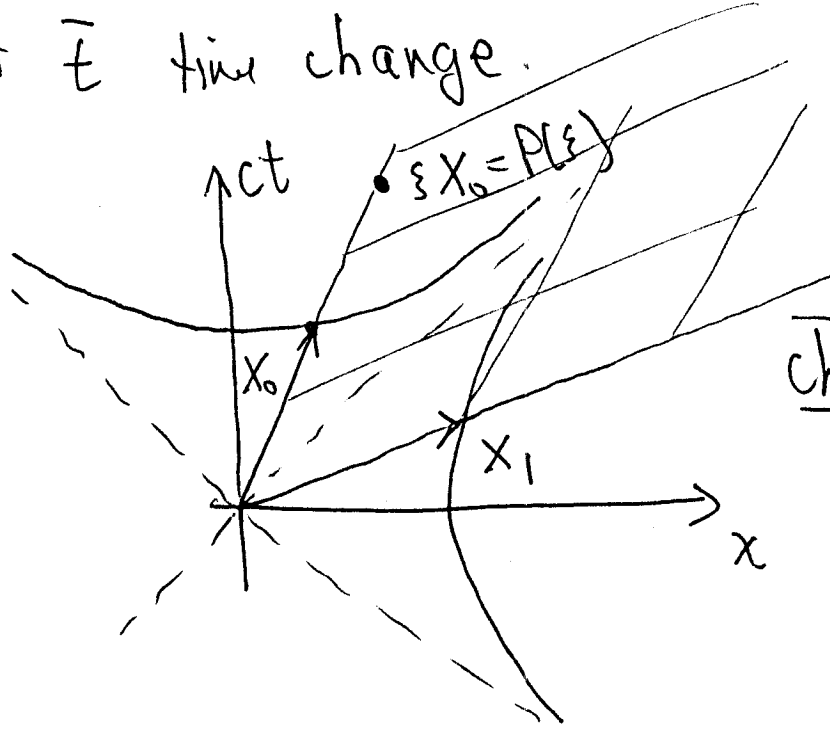
- (20)
- Time Dilation: An observer fixed in the unbarred frame (say at origin) moves along a curve $x^i = 0$, $x^0 = \xi$. His "aging time" betw t_1 and t_2 is given by

$$c\Delta\tau = \int_{ct_1}^{ct_2} \sqrt{-\eta_{ij} \dot{x}^i \dot{x}^j} d\xi = \int_{ct_1}^{ct_2} d\xi = c\Delta t$$

Conclude: Proper time & coordinate time agree for an observer fixed in L-frame:

By symmetry, an observer fixed on \bar{x}^0 -coord axis ages according to the change in his \bar{t} -coordinate.

But: starting clocks at $t = \bar{t} = 0$, the time change for observer fixed in \bar{x} -coordinates between 0 and $P(\xi) = \xi X_0$ is $\xi = \bar{t}c$ because X_0 represents a unit \bar{t} time change.



Really
 $P(\xi) = \bar{t} (\xi X_0)$

Check: $c\bar{t} = \int_0^{\xi_0} \|\xi X_0\| d\xi$
 $= \int_0^{\xi_0} \xi d\xi = \xi_0$

But in \bar{x} -coords, $\xi X_0 = (\xi \cosh \theta, \xi \sinh \theta)$

$\Rightarrow ct = \xi \cosh \theta = c\bar{t} \cosh \theta$

$\therefore \bar{t} = \frac{1}{\cosh \theta} t = \sqrt{1 - \tanh^2 \theta} t = \sqrt{1 - \left(\frac{v}{c}\right)^2} t$

$1 - \left(\frac{v}{c}\right)^2 = 1 - \tanh^2 \theta = \operatorname{sech}^2 \theta$

"Moving clocks appear to run slowly"

Conclude: $\bar{t} = \sqrt{1 - \left(\frac{v}{c}\right)^2} t$

thus $\Delta \bar{t} = \sqrt{1 - \left(\frac{v}{c}\right)^2} \Delta t < \Delta t$

"Moving observers carry clocks that appear to move slow rel to coord clocks"

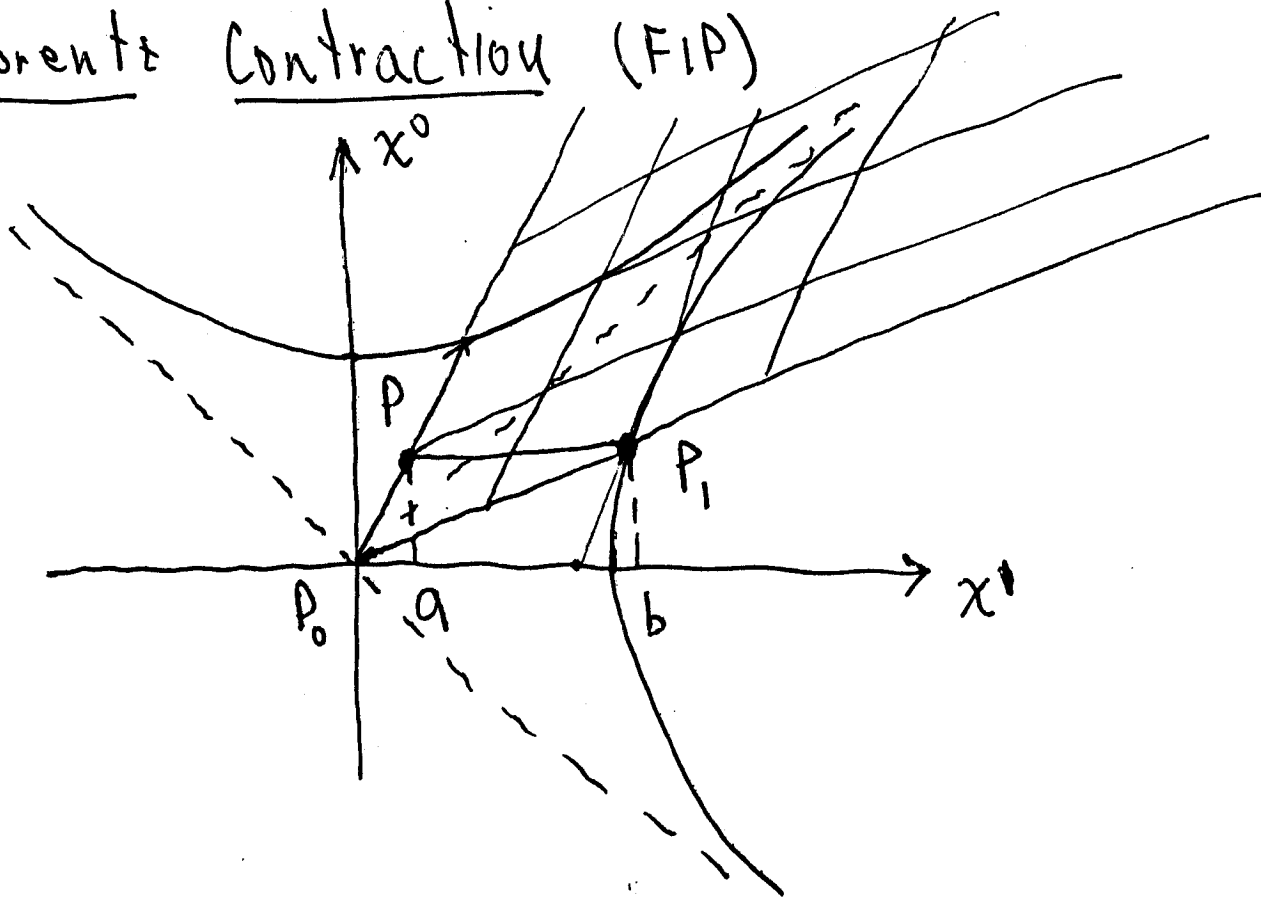
Twin Paradox: Any observer that moves & returns will age less than fixed observer.

Q: Explain lack of symmetry?

Homework: Show $L(\theta)L(\bar{\theta}) = L(\theta + \bar{\theta})$, and use this to show that if the \bar{x} -frame moves with vel v relative to x -frame, and the $\bar{\bar{x}}$ -frame moves with velocity \bar{v} rel to the barred frame, then the following velocity transformation law holds:

$$\bar{\bar{v}} = \frac{v + \bar{v}}{1 + \frac{v\bar{v}}{c^2}}$$

Lorentz Contraction (FIP)



A meter stick in the barred frame has $P_0 = 0$ and P_1 s.t. $X_1 = \frac{\partial}{\partial \bar{x}^1}$, $P_1 = \Phi(X_1)$, X_1 the spacelike unit vector in the barred $o-n$ frame. Thus $X_1 = \sinh\theta \frac{\partial}{\partial x^0} + \cosh\theta \frac{\partial}{\partial x^1} \Rightarrow P_1 = (\sinh\theta, \cosh\theta)$

$$\underline{x}(P_1) = \underline{x}_0 \Phi(X_1) = (\sinh\theta, \cosh\theta)$$

Let $L_{\underline{x}}$ denote the length of meter stick as measured in the \underline{x} -frame, (unbarred). In this frame, the length $L_{\underline{x}}$ will be

change in x' -coord. across the bar as measured at the same time x^0 . Thus, (22c)

$$L_{\underline{x}} = b - a$$

where $b = \cosh \theta$, and $a = x'$ -coord of the point P on the \bar{x}^0 -coord axis at the time $x^0 = \sinh \theta$. But \bar{x}^0 -axis in \underline{x} -coords is

$$\underline{x}(\xi) = \xi (\cosh \theta, \sinh \theta),$$

$$\text{so } x^0(\xi) = \xi \cosh \theta = \sinh \theta \Rightarrow$$

$$\xi = \tanh \theta$$

$$\Rightarrow a = \xi \sinh \theta = \tanh \theta \sinh \theta$$

$$L_{\underline{x}} = b - a = \cosh \theta - \tanh \theta \sinh \theta$$

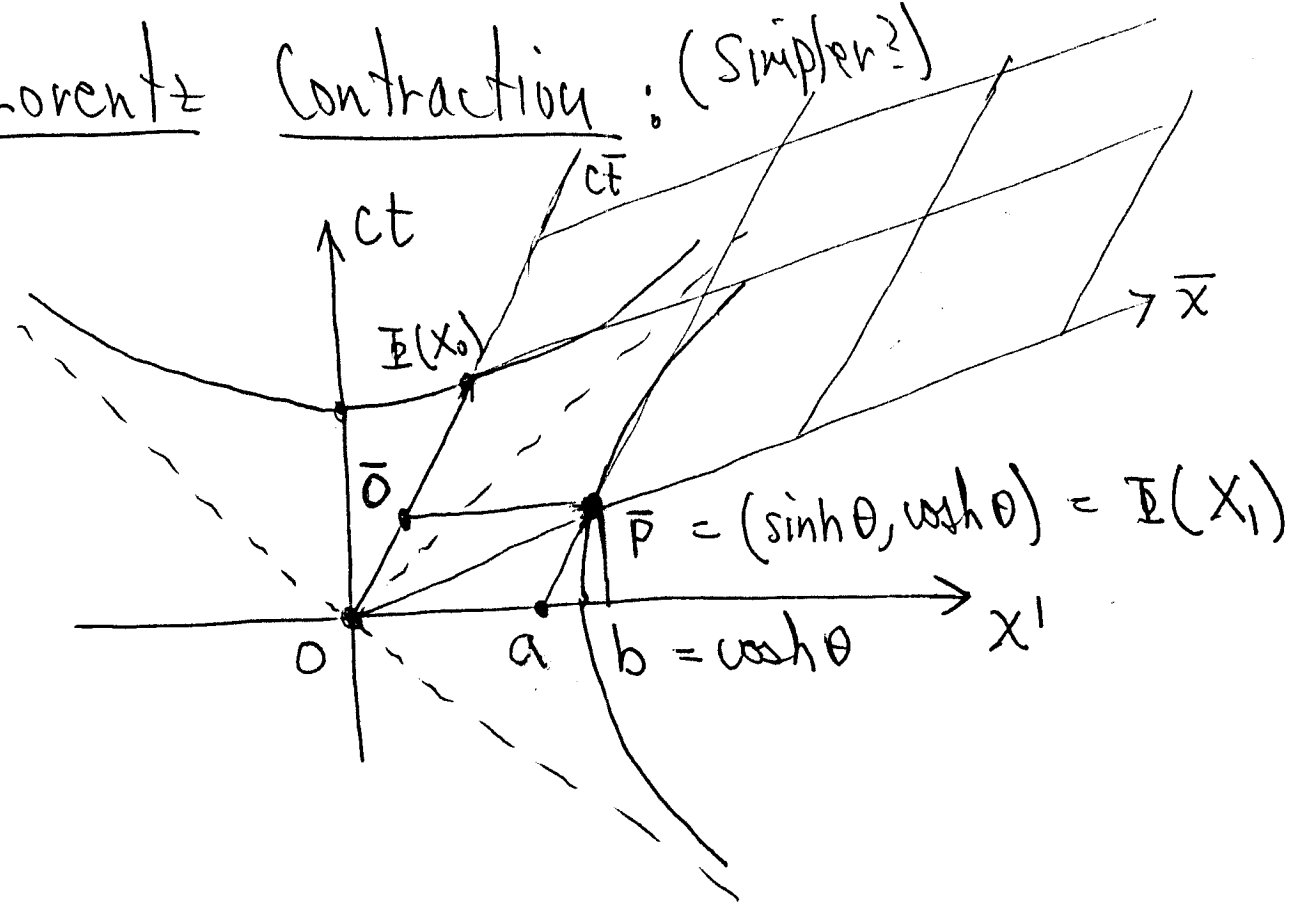
$$= \cosh \theta - \frac{\sinh^2 \theta}{\cosh \theta} = \cosh \theta - \frac{\cosh^2 \theta - 1}{\cosh \theta}$$

$$= \text{sech } \theta = \sqrt{1 - \left(\frac{v}{c}\right)^2}$$

$$L_{\underline{x}} = \sqrt{1 - \left(\frac{v}{c}\right)^2} L_{\bar{x}}$$

\Rightarrow
 "moving rods appear contracted by a factor $\sqrt{1 - \left(\frac{v}{c}\right)^2}$ "

☐ Lorentz Contraction: (Simpler?)



• The barred observer measures a unit bar fixed in his frame at $t=0$ as $\|X_1\| = 1$. The unbarred observer sees bar moving at speed v and measures bar at $t=0$ as equal to a

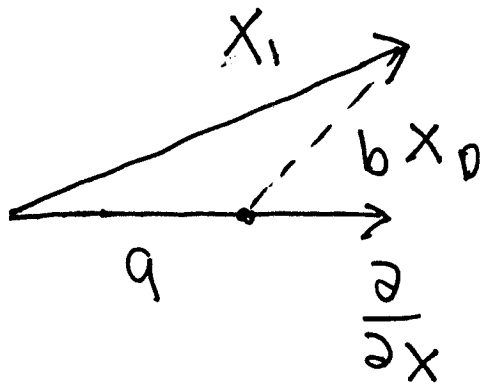
But $a = b - (b-a) = \cosh \theta - (b-a)$ and

$(b-a)$ is the x^1 -coord of $\bar{O} = \xi (\cosh \theta, \sinh \theta)$

where $\xi = \tanh \theta$ so \bar{O} and \bar{P} have same x^0 -coordinate.

Thus $a = \cosh \theta - \cosh \theta \tanh \theta = \cosh \theta \operatorname{sech}^2 \theta$
 $= \operatorname{sech} \theta = \sqrt{1 - (v/c)^2}$ ✓

• Note: For an invariant formula for the Lorentz contraction factor, note:



$$a \frac{\partial}{\partial x} + b X_0 = X_1 \equiv \text{vector along moving rod}$$

or better:

$$a \frac{\partial}{\partial x} + b X_1^\perp = X_1 \quad (X_1^\perp \perp X_1 \text{ wrt } g)$$

Thus $\langle a \frac{\partial}{\partial x} + b X_1^\perp, X_1 \rangle = \langle X_1, X_1 \rangle$

$$a \langle \frac{\partial}{\partial x}, X_1 \rangle = \langle X_1, X_1 \rangle$$

$$a^{-1} = \frac{\langle X_1, \frac{\partial}{\partial x} \rangle}{\langle X_1, X_1 \rangle} = \| \text{Proj}_{X_1} \frac{\partial}{\partial x} \|$$

Conclude: A moving rod will be measured as $\frac{1}{a}$ times longer in its rest frame than in the moving frame, where

$$\frac{1}{a} = \|\text{Proj}_{X_1} \frac{\partial}{\partial x}\| \Leftrightarrow a = \frac{1}{\|\text{Proj}_{X_1} \frac{\partial}{\partial x}\|}$$

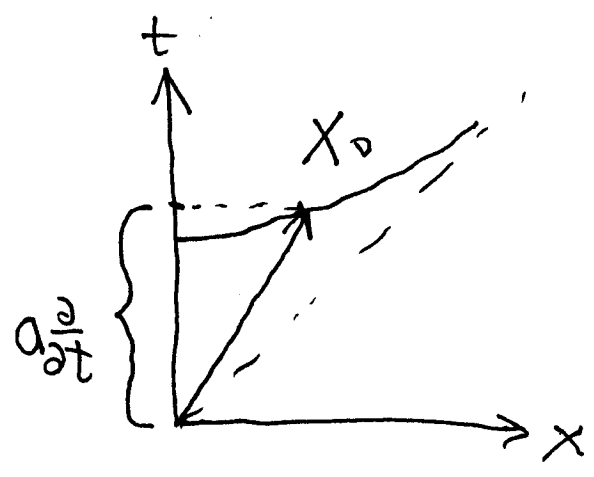
• Similarly,

$$a \frac{\partial}{\partial t} + b \frac{\partial}{\partial x} = X_0 \text{ timelike}$$

$$\langle a \frac{\partial}{\partial t}, \frac{\partial}{\partial t} \rangle = \langle X_0, \frac{\partial}{\partial t} \rangle$$

$$a = \langle \frac{\partial}{\partial t}, X_0 \rangle$$

contraction factor when measured in $\frac{\partial}{\partial x}$



= "time measured by $\frac{\partial}{\partial t}$ observer for a clock moving along X_0 "